

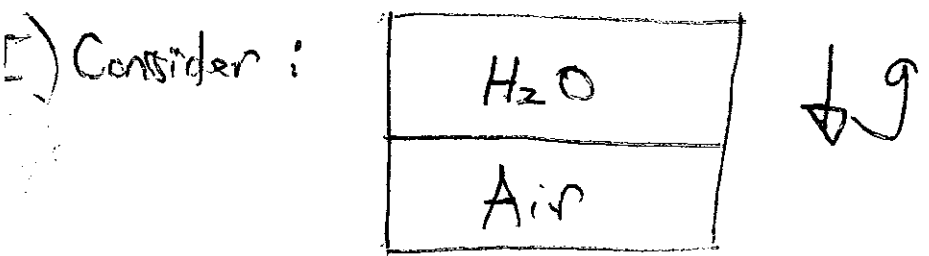
Rayleigh - Taylor Instability \leftrightarrow A Case Study

2. Motivation and ICF Overview

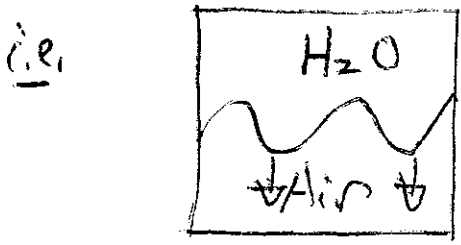
\rightarrow RT is simple example/paradigm of non-trivial nonlinear collective dynamics

\rightarrow intellectual content typical of current problems in plasma physics \rightarrow { nonlinear evolution of instabilities, turbulence, transport, etc.

Overview of RT Physics:



- free energy available (i.e. gravitational potential energy) (free energy \leftrightarrow instability) (successful storage \leftrightarrow confinement)
- system in equilibrium (i.e. inverted glass H_2O + cardboard) but small interface perturbations grow.



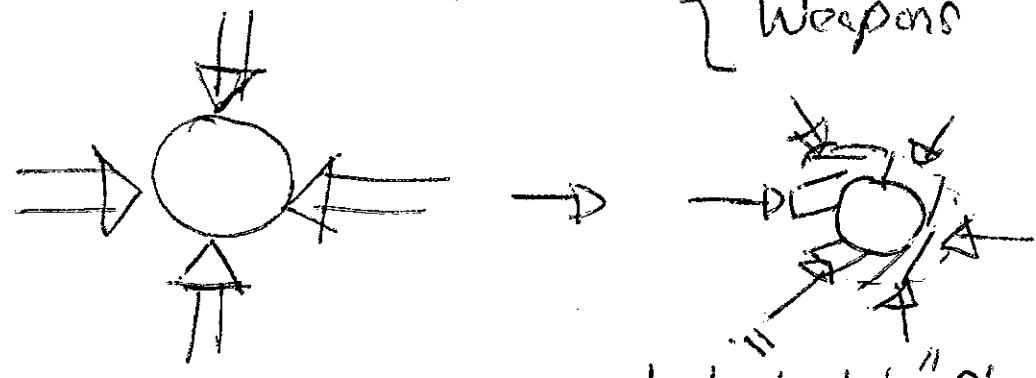
water-glass demo.

II) - typical evolutionary history:

→ instability occurs when light fluid accelerated into heavy fluid

⇔ in light fluid frame equivalent to inverted water glass

Imp: Importance R-T in ICF $\rho_1 \ll \rho_2$
 e.g. spherical implosions } ICF
 Weapons etc.



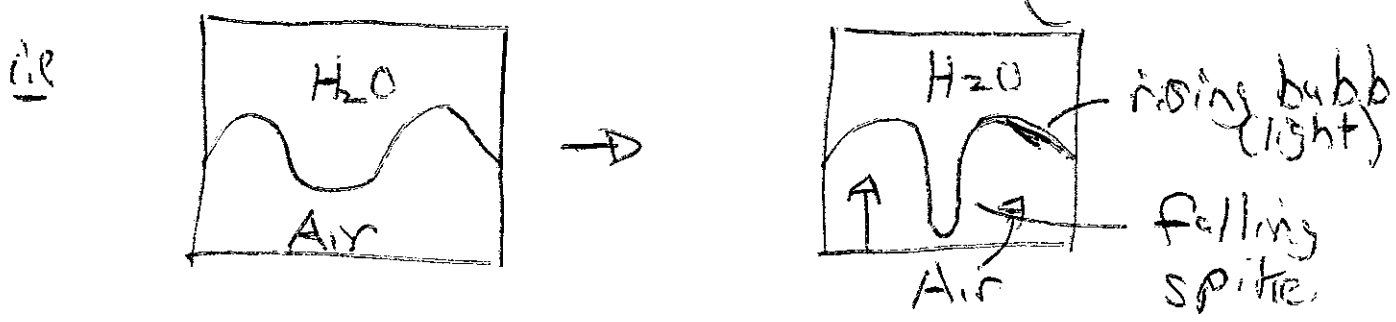
hot "light" fluid accelerated into "heavy" core
 ablation-drives rocket

① → $\epsilon < \lambda$ → linear growth phase

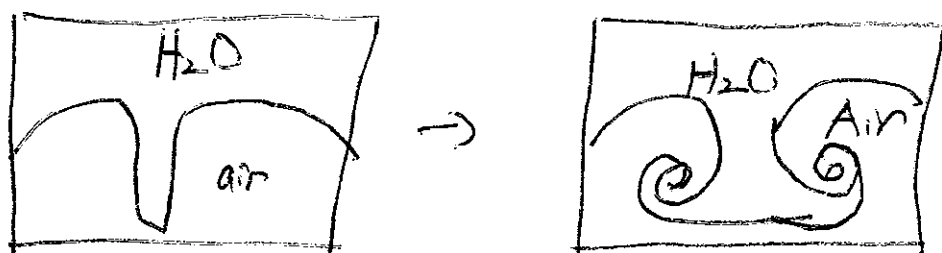
i.e. $\vec{\epsilon}_k = \vec{\epsilon}_k(0) e^{\gamma t}$

↳ calculated from linear perturbation analysis

② → $\epsilon \gtrsim \lambda$ → Spikes and Bubble Formation Competition

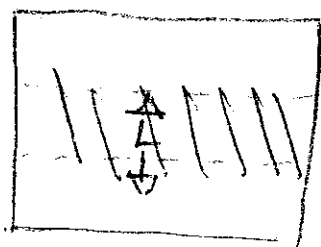


- ③ $\epsilon \gtrsim \lambda \rightarrow$ Secondary Instability / Bubble Competition
- Spike undergoes Kelvin-Helmholtz (shearing) instability
 - Spike "rolls up" and is "blunted"



- ④ $\epsilon \gg \lambda \rightarrow$ Turbulent Mix

- Spike undergoes KH \rightarrow turbulence generated
- Spike + bubble ensemble \Rightarrow mixing layer, growing in time



phenomenological

$$L \sim (0.05) \frac{(\rho_w - \rho_a) g t^2}{(\rho_w + \rho_a)}$$

intuition from elementary mech.

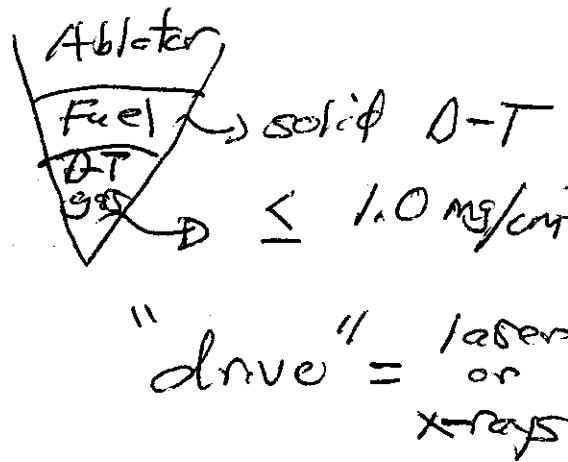
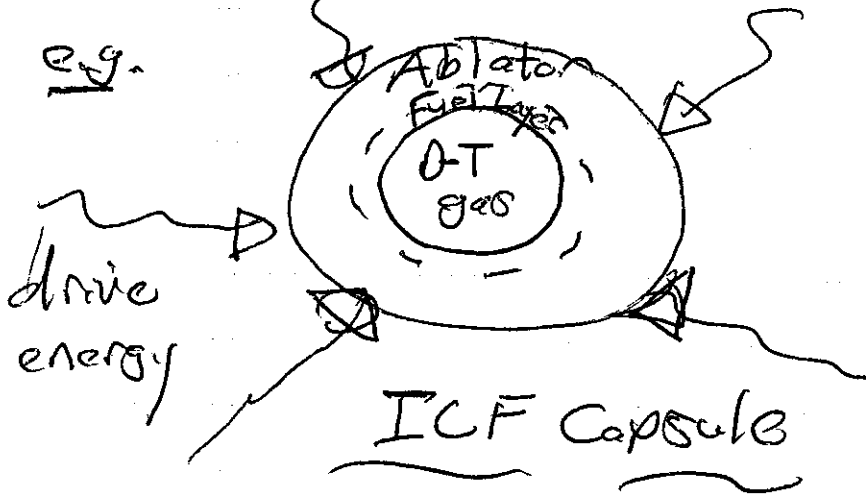
Note:

- (i) Representation
- ① \rightarrow Fourier Modes
 - ②, ③ \rightarrow Structures (Spike, Bubble)
 - ④ \rightarrow Turbulence

→ R-T. in ICF

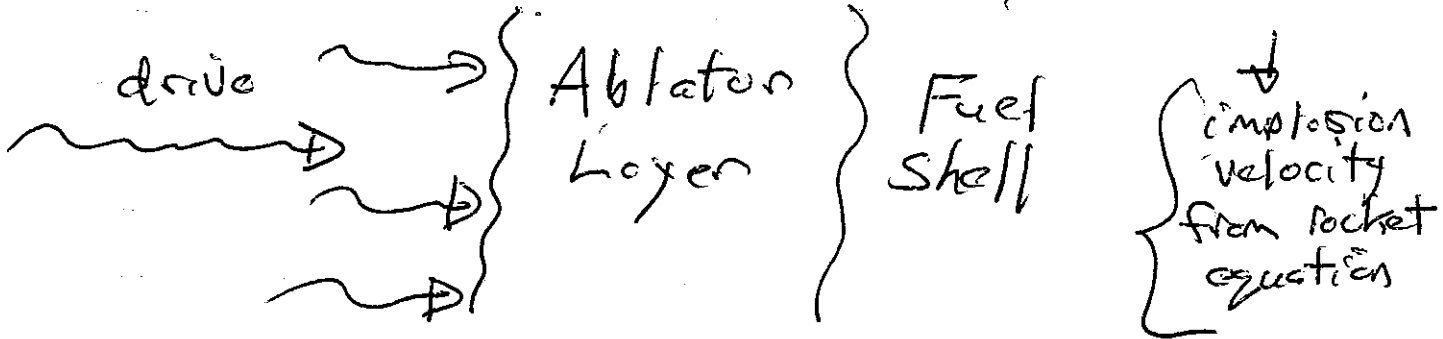
a.) Some Basics of ICF

ICF: I for Inertial

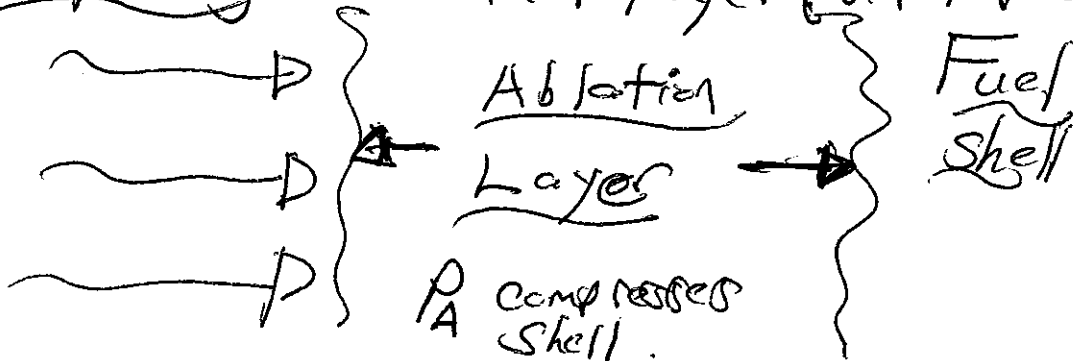


How it works:

Ablation-Driven Rocket



ablator layer heats and expands thus compressing inner fuel layer (via PV work)



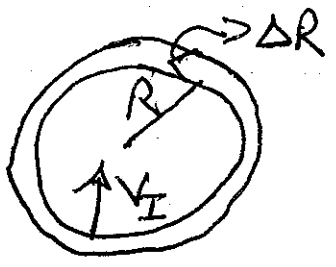
note: \rightarrow "implosion" is just conservation of momentum between expanding ablator layer and inner shell

$$\rightarrow W_{OF} \text{ (work on fuel)} \sim P_A \frac{V_s}{V_{shell}}$$

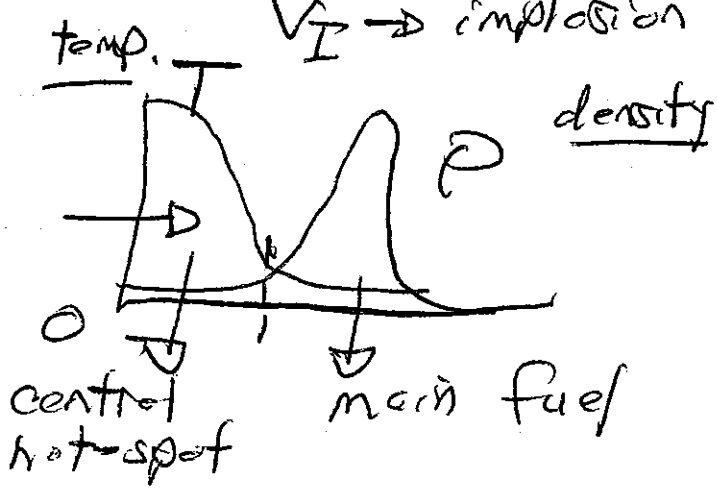
\downarrow ablation pressure

\therefore For fixed P_A (determined by driver and materials), larger, thin shells can be accelerated better than small thick ones.

\rightarrow expected (Chapal for...) final state seq:

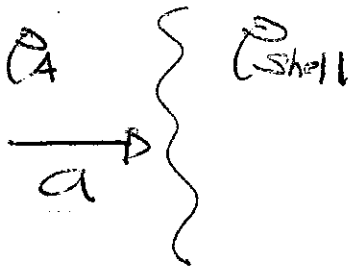


- $R \rightarrow$ shell radius
- $\Delta R \rightarrow$ shell thickness
- $V_I \rightarrow$ implosion velocity



idea is that burn initiates in central hot-spot, then propagates to main fuel shell.

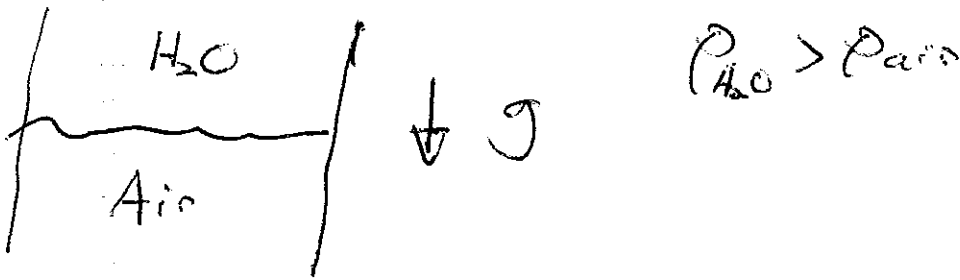
Now! Consider situation:



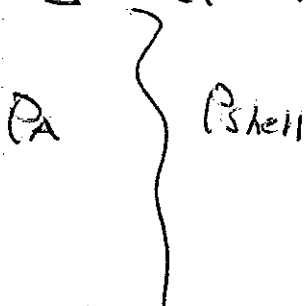
$$\rho_{shell} > \rho_A$$

i.e. Light fluid "pushing on" (i.e. accelerating into) heavy fluid

Compare to inverted glass of H_2O :



i.e. in frame of ablator, above interface:



\Rightarrow Rayleigh Taylor Instability!

\Rightarrow PGS. 1-2

Important features of Implosion :

→ IFAR - in flight aspect ratio
(→ stability)

$$IFAR = \frac{R}{\Delta R}(t)$$

$\Delta R < \Delta R(t=0)$
due compr.

→ seek large IFAR

→ but R-T_i constrains upper limit
on IFAR → broadens ΔR via mixing

i.e. $25 < IFAR < 35$

⇒ sets minimum P_A (~ 100 Mbar)
and irradiance absorbed ($\sim 10^{15}$ W/cm²)
for MJ drivers in order to achieve
 $V_I \sim 3-4 \times 10^7$ cm/sec.

⇒ R.T_i is (partly) why NIF costs
> 1 BB i.e. drives cost of laser.

→ C_r - convergence ratio
(→ symmetry)

$$C_r = R_{a,i} / r_{hot spot, f}$$

i → initial
f → final

ie deviation from sphericity can destroy hot-spot (burn-through), etc.

so

$$\delta R = \frac{1}{2} dg t^2 = \frac{dg}{g} (R_A - r) = \frac{dg}{g} r (C_n - 1)$$

\downarrow deviation from sphericity \downarrow deviation from eq. acceleration

Tolerable asymmetry ~~is~~ excess of K.E. above ignition threshold. IF demand, say

$$\delta R < \frac{r}{4} \Rightarrow \frac{dg}{g} \sim \frac{dV}{V} < \frac{1}{4(C_n - 1)}$$

since $C_n < 40$, need $\frac{dV}{V} \lesssim 1\% !!$

→ { Point is that R.T. ~~is~~ ripples ~~is~~ asymmetry, can destroy implosion via ~~is~~ inducing asymmetry, unless $K.E. \gg$ ignition threshold

\downarrow
Laser drive

once again, R.T. ~~is~~

(ii) Evolution : ① → exponential

②, ③ → transition to algebraic

④ → algebraic

III) Application ICF
Skip, in favor II. II here

Controlled Fusion $\Leftrightarrow nTt > (nTt)_{\text{Lawson}}$

Confinement → magnetic (tokamaks, etc.)

→ inertial (Laser acceleration, gravity (star))

→ ICF :

→ confine burning plasma via implosion driven by laser-produced ablation

→ implosion drives $nTt > (nTt)_{\text{Lawson}}$

Further :

→ optimal to implode shell :



acceleration → outer surface ablated
(laser pulse) → RT unstable

deceleration → inner gas accelerated into inner shell
(post pulse) → RT unstable

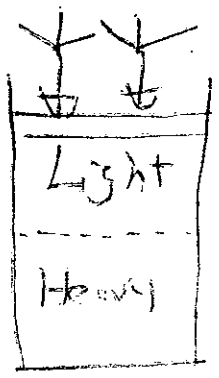
→ implosion instability intrinsic to ICF

∴ Need understand, minimize

→ Basic Insight

- Computer Simulations
- Laboratory Experiments

Experimental Set-Up (Youngs Rocket Rig, D. Youngs, AWE)



→ Rocket Engine:

- Easy:
- diagnosis
- Flow visualizations

References:

Landau, Lifshitz; Fluid Mechanics (Linear Theory)

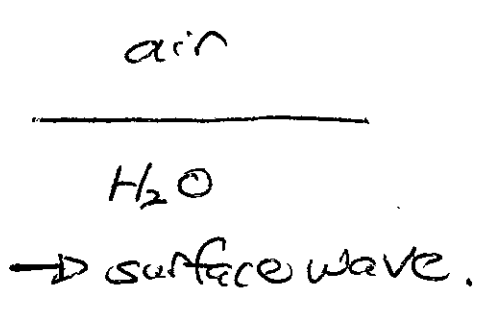
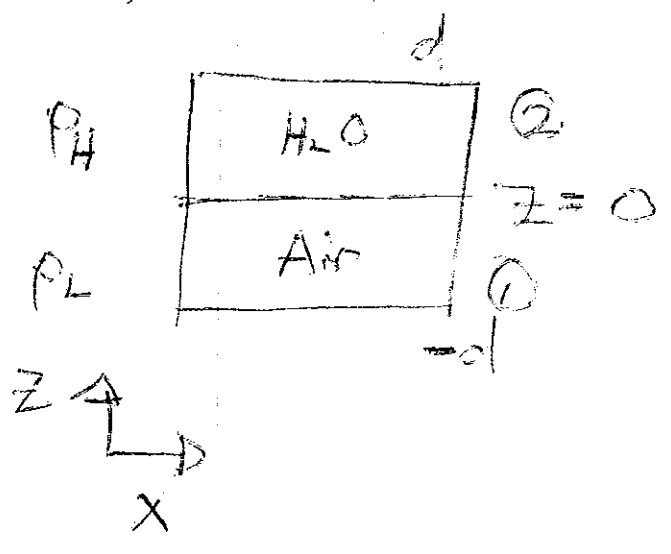
D. H. Sharp, Physics 120 (1984) p. 3 (overview)

S. Chandrasekhar "Hydrodynamic and Hydromagnetic Stability" Oxford U. Press (Linear Theory)

H. J. Kull, Physics Reports 206 #5 1991 (Review)

a.) Linear Theory

I) Hydrodynamic RT / Plane Slab



Now, consider:

- incompressible fluid (i.e. $\gamma \ll kc_s$)

$$\nabla \cdot \underline{v} = 0$$

- irrotational flow ($\nabla \times \underline{v} = \underline{\omega} = 0$)
(piecewise uniform density)

III \rightarrow Newton's tube

$$\nabla \times \underline{v} = 0 \Rightarrow \underline{v} = \nabla \phi$$

ϕ
Stream Function

$$\nabla \cdot \underline{v} = 0$$

$\Rightarrow \nabla^2 \phi = 0 \iff$ R.T. instability is potential flow problem

Now, $\phi = \sum_k \phi_k(z) e^{ikx}$ (∞ -ly wide or periodic box)

$\frac{\partial^2 \phi_k(z)}{\partial z^2} - k^2 \phi_k = 0 \Rightarrow$ origin of ρ continuity

For $kd \gg 1$, neglect finite depth, so

$\phi_k = \begin{cases} \phi_k^{(1)} e^{kz} & z < 0 \quad (1) \\ \phi_k^{(2)} e^{-kz} & z > 0 \quad (2) \end{cases}$ (satisfy $V_n = 0$ bndry)


At $z=0$:

$\rho^{(1)} = \rho^{(2)} \rightarrow$ pressure continuity (else interface motion on acoustic time scale)

$\left. \frac{\partial \phi^{(1)}}{\partial z} \right|_0 = \left. \frac{\partial \phi^{(2)}}{\partial z} \right|_0 \rightarrow$ normal velocity continuity

For dynamics:

\rightarrow described entirely by interface motion

i.e. \rightarrow 

\rightarrow fields: $\eta(x, z, t) \rightarrow$ instantaneous interface position

$\phi(x, z, t) \rightarrow$ stream fnct
 \downarrow
 $z = 0 + \eta$ (why NLT here. η dropped for linearized theory)

for stream function:

(Bernoulli's law)

$$\rho \left(\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} \right) = -\nabla p - \rho g \quad (g = -g \hat{z})$$

$$\underline{v} = \nabla \phi$$

$$\rho \left(\frac{\partial}{\partial t} \nabla \phi + \nabla \phi \cdot \nabla \nabla \phi \right) = -\nabla p - \rho g$$

$$\rho \left(\frac{\partial}{\partial t} \nabla \phi + \nabla \left(\frac{\nabla \phi \cdot \nabla \phi}{2} \right) \right) = -\nabla p - \rho g$$

$$\Rightarrow \boxed{\frac{\partial \phi}{\partial t} + \frac{\nabla \phi \cdot \nabla \phi}{2} = -\frac{p}{\rho} - g \eta} \quad (v = v_h)$$

i.e. $\frac{\partial \phi}{\partial t} = 0$ \Rightarrow $\rho + \frac{\rho v^2}{2} = \text{const.}$
 $g = 0$

For interface:

$$\boxed{\frac{\partial \eta}{\partial t} + \nabla \phi \cdot \nabla \eta = \frac{d\eta}{dt} = \frac{\partial \phi}{\partial z}} \rightarrow \text{definition}$$

Then, linearizing for R.I. mode:

$$\frac{\partial \tilde{\phi}}{\partial t} = -\frac{\tilde{p}}{\rho} - g \tilde{\eta}$$

$$\frac{\partial \tilde{\eta}}{\partial t} = \frac{\partial \tilde{\phi}}{\partial z}$$

thus:

$$\rho_2 \frac{\partial \tilde{\phi}^{(2)}}{\partial t} + g \rho_2 \tilde{\eta} = -\tilde{p}^{(2)} \quad (e^{-kz})$$

$$\rho_1 \frac{\partial \tilde{\phi}^{(1)}}{\partial t} + g \rho_1 \tilde{\eta} = -\tilde{p}^{(1)} \quad (e^{kz})$$

At interface: $\tilde{p}^{(1)}|_0 = \tilde{p}^{(2)}|_0$

$$\rho_2 \frac{\partial \tilde{\phi}^{(2)}}{\partial t} + g \rho_2 \tilde{\eta} = \rho_1 \frac{\partial \tilde{\phi}^{(1)}}{\partial t} + g \rho_1 \tilde{\eta}$$

$$\tilde{V}_z^{(1)}|_0 = \tilde{V}_z^{(2)}|_0$$

$$\Rightarrow +k \tilde{\phi}^{(1)} = -k \tilde{\phi}^{(2)}$$

\Rightarrow

$$g(\rho_2 - \rho_1) \tilde{\eta} = \rho_1 \frac{\partial \tilde{\phi}^{(1)}}{\partial t} - \rho_2 \frac{\partial \tilde{\phi}^{(2)}}{\partial t}$$

$$= (\rho_1 + \rho_2) \frac{\partial \tilde{\phi}^{(1)}}{\partial t}$$

$$\frac{\partial \tilde{\phi}^{(1)}}{\partial t} = \frac{g(\rho_2 - \rho_1)}{(\rho_1 + \rho_2)} \tilde{\eta}$$

$$\frac{\partial \tilde{\eta}}{\partial t} = \frac{\partial \tilde{\phi}^{(1)}}{\partial z}$$

$$\frac{\partial^3 \tilde{\phi}^{(1)}}{\partial z^3} = g \frac{(\rho_2 - \rho_1)}{(\rho_2 + \rho_1)} \frac{\partial \tilde{\phi}^{(1)}}{\partial z}$$

⇒

$$\omega_1^2 = -g A k$$

$$\boxed{\gamma = \sqrt{g A} \sqrt{k}}$$

$A \equiv \frac{\rho_2 - \rho_1}{\rho_2 + \rho_1}$
 Atwood # - available free energy

Comments:

(i) equivalent: { fluid with vacuum } $\rho \rightarrow A$

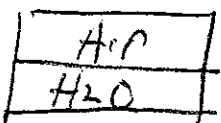
(ii) H₂O, air: $\lambda = 1 \text{ cm}$ $\gamma \sim 1 \text{ sec}^{-1}$
 (fast)

(iii) $\gamma = \sqrt{g A k}$

∴ in absence of dispn, surface tension etc,
 shorter wavelengths grow faster

(iv) $A < 0 \Rightarrow$ stable stratification
 → surface buoyancy wave

H₂O, Air $\Rightarrow \omega = \sqrt{k g}$ → surface gravity wave



- Other Effects:

(i) Surface Tension (Fluid) \rightarrow III (HW)

- curvature of interface exerts force

i.e. $\rho \rightarrow \rho - \rho \gamma_T \nabla_h^2 \eta$ ($\gamma_T = \frac{T_s}{\rho}$)

(For H₂O - air, only H₂O feels surface tension; for fluid ①, fluid ②, T_s for each interface)

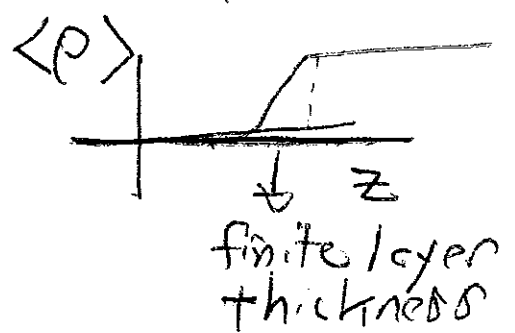
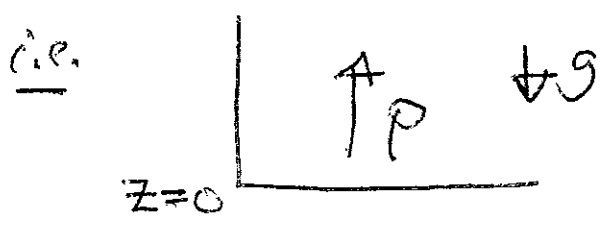
$\Rightarrow \gamma = (kgA - \gamma_T k^3)^{1/2}$

$k_{max} = (gA / \gamma_T)^{1/2}$
unstable

\rightarrow range of modes limited

eg. inverted glass with cardboard $\rightarrow \gamma_T \rightarrow \infty$

(ii) Finite Interface Thickness - $\nabla \rho$



Consider opposite limit:

$k L_p \gg 1$

$L_p = \frac{1}{\rho_0} \frac{d\rho}{dz}$

rippled interface \rightarrow cell

- fluid motion not irrotational, as $\nabla \rho \neq \nabla p$
hydrostatic eqn $\frac{d\rho}{dz} = -\rho g$

Review

→ Last time:

$$\nabla^2 \phi = 0$$

$$\Rightarrow \begin{cases} \phi_H = \tilde{\phi}_H e^{-kz} \\ \phi_L = \tilde{\phi}_L e^{kz} \end{cases}$$

$$\begin{bmatrix} \rho_H \\ \rho_L \end{bmatrix}$$

$$\begin{cases} \frac{\partial \phi}{\partial t} + \frac{(\nabla \phi)^2}{2} = -\frac{p}{\rho} - g\eta \quad \rightarrow \text{Bernoulli} \\ \frac{\partial \eta}{\partial t} + \nabla \phi \cdot \nabla \eta = \frac{\partial \phi}{\partial z} \quad \rightarrow \text{defn.} \end{cases}$$

$$\tilde{v}_{Hz} = \tilde{v}_{Lz} \quad \tilde{\rho}_H = \tilde{\rho}_L$$

L.T. \Rightarrow

$$\gamma = \sqrt{gAk}$$

$$A = \left[\frac{\rho_H - \rho_L}{\rho_H + \rho_L} \right]$$

→ Key Assumptions:

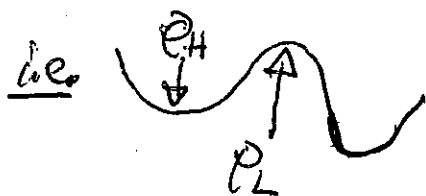
- incompressible $\rightarrow \gamma \ll k c_s$
- inviscid $\rightarrow \gamma \gg \nu k^2$

- irrotational $\rightarrow \underline{v} = \nabla \phi$

- this interface ("piecewise uniform")
- no breaking \rightarrow amplitude restricted
- no k.H.

$\Delta \rightarrow$ potential flow.

interface ripples



but "heavy" falls
"light" rises

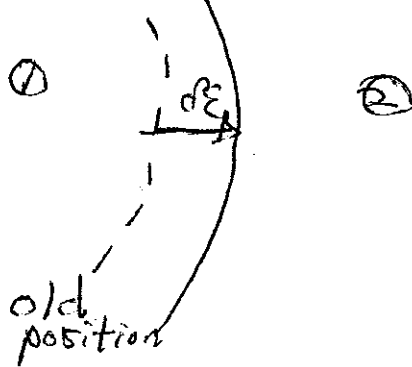
Insert ~~III~~ Surface Tension

→ Consider two liquids separated by a thin (i.e. few molecules) interface



Now, consider displacing the interface toward ② by $d\epsilon$

i.e.



($\epsilon \leftrightarrow \Delta$)

∴ can determine change in free energy (i.e. thermodynamic sense) via:

$$dF = \underbrace{dF_1 + dF_2}_{\text{bulk phases}} + dF_{\text{interface}}$$

↳ treat as separate constituents

Recall: $dF = -SdT - pdV$

(i.e. $F = E - ST$)

$$dF_{1,2} = (-SdT - pdV)_{1,2}$$

(i.e. the usual)

then, for interface, natural to define :

$$dF_I = -S_I dT + \underline{\sigma} dA,$$

↓
entropy
of interface

↳ change in free energy
due increase in surface

area of interface (treat as
separate phase)

$\sigma \equiv$ Surface Tension

(\sim Pressure \times Length \sim Force/Length
(interf. σ /Area \times L))

Hereafter, consider isothermal displacement.

$$\rightarrow dF = -p_1 dV - p_2 (-dV) + \sigma dA$$

$$= (p_2 - p_1) dV + \sigma dA$$

interface expands 'into' 2nd material

Further: $dV = dA d\varepsilon$ (for surface)

↓
displacement

→ $\varepsilon(x,y)$

$$\text{For } dA: \quad dA = \int dx dy \left(1 + \left(\frac{\partial \varepsilon}{\partial x} \right)^2 + \left(\frac{\partial \varepsilon}{\partial y} \right)^2 \right)^{1/2} - \int dx dy$$

∴ for small displacement:

$$dA \approx \int dx dy \left(1 + \frac{1}{2} \left(\frac{\partial \varepsilon}{\partial x} \right)^2 + \frac{1}{2} \left(\frac{\partial \varepsilon}{\partial y} \right)^2 \right)$$

$$\therefore dA = \int dx dy (-\nabla^2 \Sigma) d\Sigma$$

\downarrow
 curvature of
 surface displacement

(i.e. anticipates
integration by parts)

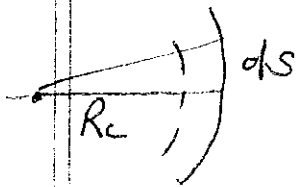
$$\underline{\text{so}} \quad dF = \int \left[(\rho_2 - \rho_1) dA_0 - \nabla^2 \Sigma dA_0 \right] d^3 \Sigma$$

\Rightarrow condition for equilibrium:

$$\rho_2 - \rho_1 = \nabla^2 \Sigma(x, y)$$

N_1 also generally: $dF = (\rho_2 - \rho_1) dA_0 d\Sigma + \nabla dA$

Now consider arbitrary (i.e. not weakly curved interface)



$$ds' = (R_c + d\Sigma) d\theta$$

$$= dl_0 \left(1 + \frac{d\Sigma}{R_1} \right)$$

In general, surface parametrized by 2 radii of curvature R_1, R_2

$$\underline{\text{so}} \quad dA = \int dl_1 dl_2 \left(1 + \frac{d\Sigma}{R_1} \right) \left(1 + \frac{d\Sigma}{R_2} \right) - \int dl_1 dl_2$$

$$dA = \int h_1 dh_2 \left(\frac{1}{R_1} + \frac{1}{R_2} \right) d\epsilon$$

Thus, have most general expression:

$$dF = \int \left[(p_2 - p_1) dA_0 - \left(\frac{1}{R_1} + \frac{1}{R_2} \right) dA_0 \right] d\epsilon$$

thus, for equilibrium with interface:

$\nabla \left(\frac{1}{R_1} + \frac{1}{R_2} \right) = -(p_2 - p_1)$	Laplace's Law
--	---------------

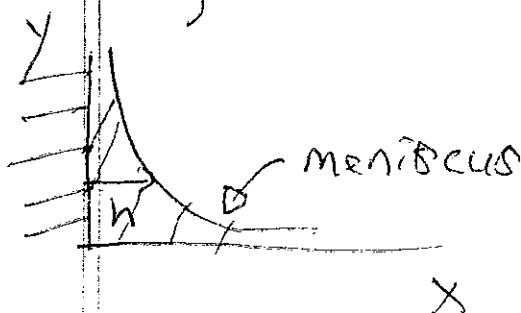
i.e. \rightarrow given 2-phase equilibrium (separated), can use to estimate droplet size for immiscible liquids

i.e. if $p_2 < p_1$

\therefore droplets of size $R \sim \sigma / (p_1 - p_2)$ may be expected.

\downarrow skip to IS

\rightarrow consider liquid adjacent to fixed vertical wall, then:



$h(y) \equiv$ defined thickness of meniscus

Then, can write:

→ known

$$p_{\text{free}} = p_0 - \rho g y(x) \quad (g < 0)$$

to calculate $h(y)$, use Laplace's Law:

c.e.
$$p_0 - \rho g y = \frac{\sigma}{R_c}$$

but
$$\frac{1}{R_c} = \frac{-\partial^2 h / \partial x^2}{(1 + (\partial h / \partial x)^2)^{3/2}}$$

c.e. don't make small curvature approx)

then taking $y_0 = 0$ (ref):

$$+ \rho g y(x) = + \frac{\partial^2 h(y) / \partial x^2}{(1 + (\partial h / \partial x)^2)^{3/2}}$$

and can get dh/dy , etc.

*
→ Capillary Waves.

Recall discussed ocean waves (stable R.T.)



Should be apparent now that:

→ for high k , curvature of crests, etc. becomes sharp

→ before, tacitly took $\rho g \eta \gg \frac{\sigma}{R_c}$
 now if $R_c \sim \eta$ s/t $\eta^2 \sim \sigma / \rho g$

must retain surface tension in ~~the~~
 surface wave dynamics \Rightarrow capillary waves

To include:

$$\rho = \rho_0 - \sigma \nabla^2 \eta$$

Then recall: $\frac{\partial \tilde{\phi}}{\partial t} = -\frac{\tilde{p}}{\rho} - g \tilde{\eta}$

$$\frac{\partial \tilde{\eta}}{\partial t} = \frac{\partial \tilde{\phi}}{\partial z}$$

$$\therefore \frac{\partial \tilde{\phi}}{\partial t} = \frac{\sigma}{\rho} \nabla^2 \tilde{\eta} - g \tilde{\eta}$$

$$\frac{\partial \tilde{\eta}}{\partial t} = \frac{\partial \tilde{\phi}}{\partial z}$$

$$\frac{\partial^2 \phi}{\partial t^2} = \frac{\sigma}{\rho} \nabla^2 \frac{\partial \phi}{\partial z} - g \frac{\partial \phi}{\partial z}$$

$$\Rightarrow \boxed{\omega^2 = kg + \frac{\sigma k^3}{\rho}} \quad \Rightarrow \text{dispersion relation for capillary waves}$$

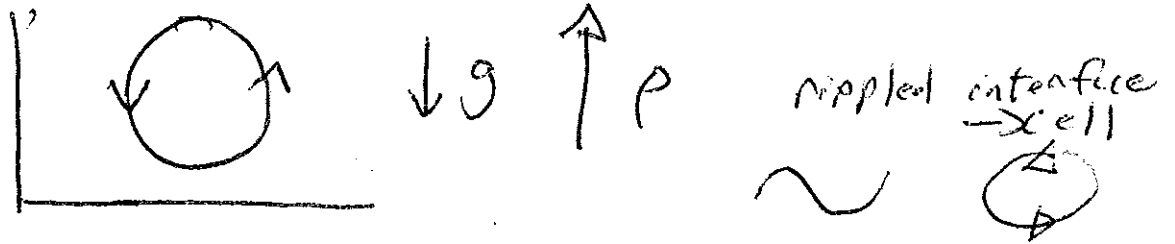
note: - capillarity estimate $\propto \sigma/\rho g$

$$\text{d.r.} \Rightarrow k_{\text{cap}}^2 \sim (\rho g / \sigma)$$

- in ocean, capillarity significant at ≤ 50
- if R.T. unstable, capillarity will cut-off high k instability

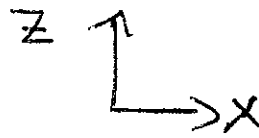
$$\text{i.e.} \quad \omega^2 = \frac{-kg(\rho_2 - \rho_1)}{(\rho_2 + \rho_1)} + \frac{\sigma k^3}{\rho_2 + \rho_1}$$

- motion is that of convective cells, vortices



To calculate:

- For 2D cell



$$\frac{\partial \tilde{v}_x}{\partial t} = -\partial_x \left(\frac{p}{\rho_0} \right)$$

$$\frac{\partial \tilde{\rho}}{\partial t} = -\tilde{v}_z \frac{d\rho_0}{dz}$$

$$\frac{\partial \tilde{v}_z}{\partial t} = -\partial_z \left(\frac{p}{\rho_0} \right) - g \frac{\partial \tilde{\rho}}{\partial z}$$

Suggests write:

$$\underline{v} = \underline{\nabla} \phi \times \underline{\hat{y}}$$

$$\Rightarrow \tilde{v}_x = -\partial_z \tilde{\phi}$$

$$\tilde{v}_z = \partial_x \tilde{\phi}$$

$$-\frac{\partial}{\partial t} \partial_z \tilde{\phi} = -\partial_x \left(\frac{p}{\rho_0} \right) \quad (1)$$

$$+\frac{\partial}{\partial t} (\partial_x \tilde{\phi}) = -\partial_z \left(\frac{p}{\rho_0} \right) - g \frac{\tilde{\rho}}{\rho_0} \quad (2)$$

$$\partial_z (1) - \partial_x (2) \Rightarrow$$

$$-\frac{\partial}{\partial t} \nabla^2 \tilde{\phi} = \frac{\partial}{\partial x} \left(g \frac{\tilde{\rho}}{\rho_0} \right)$$

$-\nabla^2 \phi = \omega_y$
 \downarrow
 \hat{y} component
 vorticity

$$\frac{\partial}{\partial t} \nabla^2 \tilde{\phi} = -\frac{\partial}{\partial x} (g \tilde{\rho} / \rho_0)$$

$$\frac{\partial}{\partial t} \tilde{\rho} = -\partial_x \tilde{\phi} \frac{d\rho_0}{dz}$$

$$\frac{\partial^2}{\partial t^2} \nabla^2 \tilde{\phi} = \left(\frac{g}{\rho_0} \frac{d\rho_0}{dz} \right) \frac{\partial^2 \tilde{\phi}}{\partial x^2}$$

$$\Rightarrow +\omega^2 k^2 = \left(\frac{g}{\rho_0} \frac{d\rho_0}{dz} \right) (-k_x^2)$$

$$\omega^2 = -\frac{k_x^2}{k^2} \left(\frac{g}{\rho_0} \frac{d\rho_0}{dz} \right)$$

$\hookrightarrow > 0$, as $d\rho_0/dz > 0$

$$\gamma = \sqrt{\frac{k_x^2}{k^2} \left(\frac{g}{L_p} \right)^{1/2}} \rightarrow \text{R.T. Convective cell growth-rate}$$

Then:

→ Structure similar to Rayleigh - Benard convection

i.e. $\frac{\partial}{\partial t}$ vorticity = torque $\left\{ \begin{array}{l} \text{buoyancy (RB)} \\ \text{gravitational force (RT)} \end{array} \right.$

$\rightarrow k_x \rightarrow \infty \Rightarrow \gamma \rightarrow \frac{g}{L_0}$

Thus, to incorporate finite interface thickness in RT growth formula

$$\begin{aligned} \gamma &\sim \sqrt{g A k} & k L_p < 1 \\ &\sim \sqrt{g/L_p} & k L_p > 1 \end{aligned}$$

$\Rightarrow \gamma = (g A k / (1 + k L))^{1/2}$

scale factor, interface.

$\therefore k L > 1 \Rightarrow$ growth rate saturates!

\rightarrow For stable stratification $d\rho_0/dz < 0$

$$\omega^2 = \frac{k_x^2}{k^2} \frac{g}{\rho_0} \left| \frac{d\rho_0}{dz} \right| \equiv \frac{k_x^2}{k^2} N^2 \rightarrow \text{BV freq}$$

\rightarrow dispersion relation for oceanic internal wave

\rightarrow finite density gradient analogue of (interface) surface wave

\sim interesting to note effects of

viscosity
particle diffusivity

viscosity $\frac{\partial}{\partial t} \nabla^2 \phi \rightarrow \left(\frac{\partial}{\partial t} - \nu \nabla^2 \right) \nabla^2 \phi$

diffusivity $\frac{\partial}{\partial t} \rho \rightarrow \left(\frac{\partial}{\partial t} - D \nabla^2 \right) \rho$

$\Rightarrow (\omega + i\nu k^2)(\omega + iDk^2) = -\frac{k_x^2}{k^2} \frac{g}{\rho_0} \frac{d\rho_0}{dz}$

i.e.
 $\left\{ \begin{array}{l} \nu k^2 \gg \omega \\ D \rightarrow 0 \end{array} \right.$ (viscous fluid)

$(i\nu k^2)(i\nu) = -\frac{k_x^2}{k^2} \frac{g}{\rho_0} \frac{d\rho_0}{dz}$

$\gamma = \frac{k_x^2}{k^2} \left(\frac{g}{\rho_0} \frac{d\rho_0}{dz} \right) / \nu k^2$

$\rightarrow \gamma \sim 1/\nu k^2$

\rightarrow strong viscosity reduces growth rate but instability persists (i.e. molasses + air!)

i.e.

$\Rightarrow D = \nu$

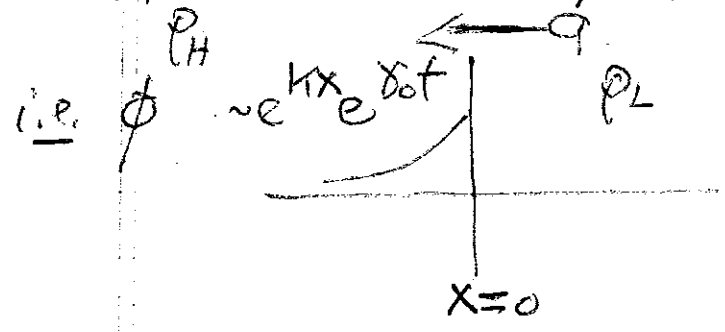
$\gamma^{\bullet} = \left(\frac{k_x^2}{k^2} \frac{g}{\rho_0} \frac{d\rho_0}{dz} \right)^{1/2} - \nu k^2$

i.e. viscosity and diffusivity can stabilize
RT instability
→ defines critical $\Delta \rho$

i.) Ablation (Ablation critical element of environment implosion ~~driven~~ ablation driven rocket)

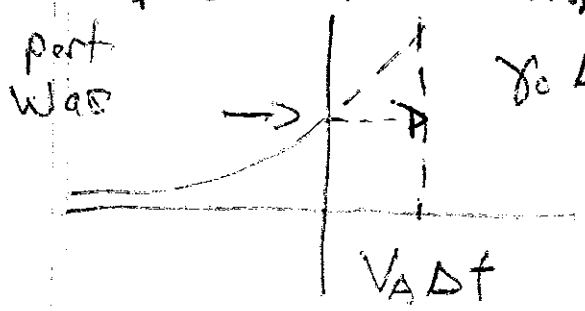
→ physical concept is that due to heating, material streams away from interface
 ∴ can't participate in RT instability

→ heuristic interpretation:



e^{+kx} ($x < 0$) is wave-function shape.

with ablation, hot matter "blown off"
 ⇒ interface displaced inward



$\gamma_0 \Delta t - k V_A \Delta t$

i.e. $e^{k(x - V_A \Delta t)}$
 ↓ re-absorb

burn-off ⇒ interface moves inward

i.e. $\delta \phi \sim e^{k(x - V_A \Delta t)} e^{\gamma_0 \Delta t}$
 $\sim e^{kx} e^{(\gamma_0 - k V_A) \Delta t}$

$V_{Ab1} \equiv \frac{M}{\rho A A}$

∴ ablative blow-off yields stabilizing effect

$\gamma = \gamma_0 - k V_A$; $\gamma_0 = \sqrt{k g}$

→ no simple, rigorous analytical theory exists!

Aside: For ICF, can combine finite interface thickness and ablative stabilization to control RT growth (A=1)

ie. simple RT $\gamma = \sqrt{k g}$

finite interface $\rightarrow \gamma = \left(k g / (1 + k L_0) \right)^{1/2}$

ablation $\rightarrow \gamma = \left(k g / (1 + k L_0) \right)^{1/2} - k V_A$

By $\left. \begin{array}{l} - \text{target design } - L_0 \\ \text{(structure)} \\ - \text{materials, etc } - V_A \\ \text{(doping)} \end{array} \right\} \text{ can minimize implosion pert. growth}$

(V_o) Spherical Geometry - Postpone till later

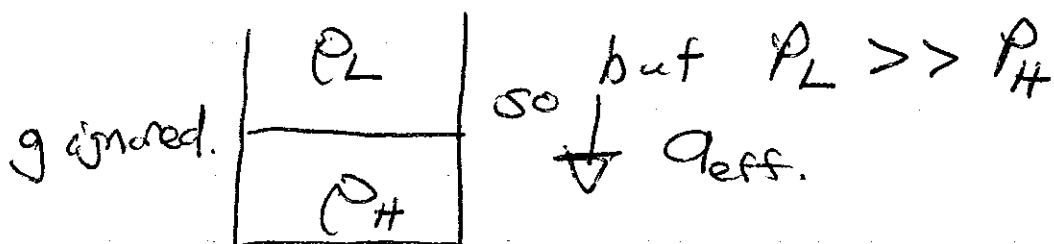
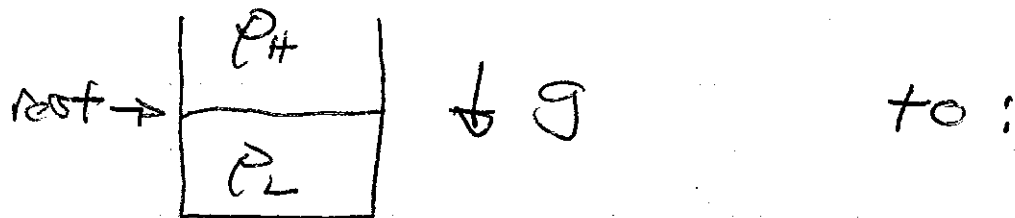
Crudely:
$$\begin{cases} \omega \sim \sigma/u \\ Lu \sim \# \sqrt{g\lambda} \end{cases}$$

N.B. $\left\{ \begin{array}{l} \text{Can solve 3 bubble Layer} \\ \text{model (numerically) to determine } \# \\ \text{in merger rule.} \end{array} \right.$

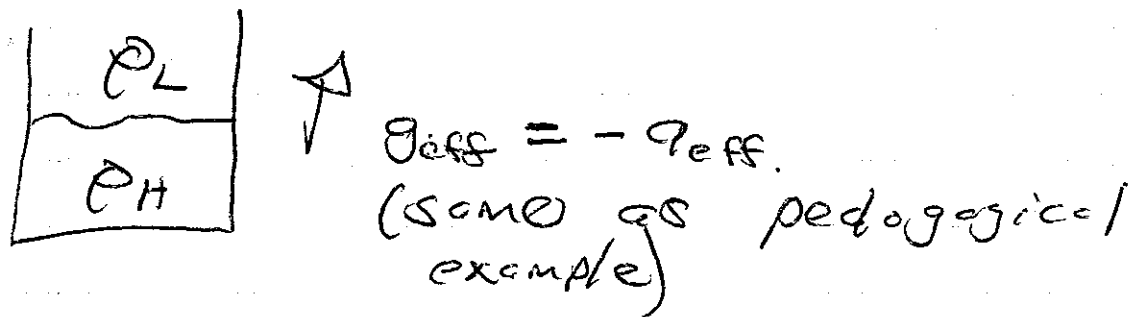
→ HW on Tuesday
 → no note borrowing on Tues, Th. AM.

19a.

Note equivalence:



in frame of interface/membrane,
 equivalent to



In general:

"R-T" occurs when light fluid
 'accelerated into/toward' heavy fluid.